#### **ORIGINAL PAPER**



# Rotational spectrum simulations of asymmetric tops in an astrochemical context

Julia C. Santos<sup>1</sup> · Alexandre B. Rocha<sup>1</sup> · Ricardo R. Oliveira<sup>1</sup>

Received: 5 June 2020 / Accepted: 26 August 2020 / Published online: 22 September 2020 © Springer-Verlag GmbH Germany, part of Springer Nature 2020

#### Abstract

Rotational spectroscopy plays a major role in the field of observational astrochemistry, enabling the detection of more than 200 species including a plethora of complex organic molecules in different space environments. Those line detections allow correctly determining the sources and physical properties, as well as exploring their morphology, evolutionary stage, and chemical evolution pathways. In this context, quantum chemistry is a powerful tool to the investigation of the molecular inventory of astrophysical environments, guiding laboratory experiments and assisting in both line assignments and extrapolation of the experimental data to unexplored frequency ranges. In the present work, we start by briefly reviewing the rotational model Hamiltonian for asymmetric tops beyond the rigid-rotor approximation, including rotational-vibrational, centrifugal, and anharmonic effects. Then, aiming at further contributing to the recording and analysis of laboratory microwave spectroscopy by means of accessible, less demanding quantum chemical methods, we performed density functional theory (DFT) calculations of the spectroscopic parameters of astrochemically relevant species, followed by their rotational spectrum simulations. Furthermore, dispersion-correction effects combined with different functionals were also investigated. Case studies are the asymmetric tops H<sub>2</sub>CO, H<sub>2</sub>CS, c-HCOOH, t-HCOOH, and HNCO. Spectroscopic parameter predictions were overall very close to experiment, with mean percentage errors smaller than 1% for zeroth order and  $\sim 5\%$  for first-order constants. We discuss the implications and impacts of those constants on spectrum simulations, and compare line-frequency predictions at millimeter wavelengths. Moreover, theoretical spectroscopic parameters of c-HCOOH and HNCO are introduced for the first time in this work.

Keywords Microwave spectroscopy · Astrochemistry · Asymmetric tops · Rotational spectroscopy

#### Introduction

A conspicuous attribute of rotational spectroscopy consists in its capability to provide a myriad of information on the

**Electronic supplementary material** The online version of this article (https://doi.org/10.1007/s00894-020-04523-0) contains supplementary material, which is available to authorized users.

This paper belongs to Topical Collection XX-Brazilian Symposium of Theoretical Chemistry (SBQT2019)

Alexandre B. Rocha rocha@iq.ufrj.br

Ricardo R. Oliveira rrodrigues.iq@gmail.com molecular species and its environment. Accurate gas-phase structure determinations can be obtained through the analysis of rotationally resolved molecular spectra [1], which provide insight on the species' fine and hyperfine structures by the means of high-resolution techniques such as chirped-pulse [2] and Lamb-dip spectroscopy [3, 4]. Moreover, pioneer laboratory detections of small molecules have been consistently achieved through rotational spectroscopy in the ranges of millimeter and sub-millimeter wavelengths [5–9].

In the field of observational astrochemistry, rotational spectroscopy is majorly employed to unveil the chemical composition and evolution of astrophysical environments. Specially, the interstellar medium (ISM) is known for its chemical lavishness, with radiofrequency line detections of more than 200 species including small hydrocarbons, molecular radicals, and a plethora of complex organic molecules (COMs) to date [10–17]. The analysis of rotational spectra also allows the investigation of the

<sup>&</sup>lt;sup>1</sup> Instituto de Química, Universidade Federal do Rio de Janeiro, UFRJ, Av. Athos da Silveira Ramos, 149, Rio de Janeiro, RJ 21941-909, Brazil

morphology and evolutionary stage of a particular environment or body [18–20], as well as the determination of its physical properties [21, 22]. Recent technological advances on single dish and interferometry radio antennas have resulted in facilities with remarkable sensitivity and resolution (e.g., the Atacama Large Millimeter Array), consequently increasing the demand for more precise rotational spectroscopic data.

In this context, quantum chemistry is a powerful tool for the investigation of molecular inventory in astrophysical environments. Theoretical predictions of spectroscopic parameters can guide laboratory experiments, assisting in both line assignments and extrapolation of the experimental data to unexplored frequency ranges [5, 9, 11, 23–34]. Moreover, the correct determination of molecular properties such as equilibrium structures from rotational spectra relies greatly on anharmonic vibrational corrections and vibrational-rotational coupling [35]. Quantum-chemical calculations at different levels of theory are therefore of great relevance to contemporary rotational spectroscopy, resulting in a notable increase in published works on the matter vis-à-vis the last century [24, 27, 31, 32, 36–39].

In this work, we focus on a theoretical approach to simulate rotational spectra of asymmetric tops (i.e., rotational constants  $A \neq B \neq C$ ) that correspond to the vast majority of molecules. The reader is referred to specialized literature on other types of rotors (e.g., [40-42] and references therein), which present much simpler energy level structures and consequently much less costly simulations. Although state-of-the-art methods which include complete basis set extrapolations and perturbative quadruple excitation in the coupled-cluster expansion (see Puzzarini et al. [43] and references therein) typically result in very accurate parameters, they are also substantially time-consuming. Consequently, more accessible approaches are beneficial in many different scenarios, Density Functional Theory (DFT) calculations of spectroscopic parameters for four different systems followed by rotational spectrum simulations were performed, aiming at further contributing to accurate and effectively guiding laboratory experiments and, consequently, facilitating the assignment and interpretation of radio spectra from the ISM. For the case studies, the astrochemically relevant species H<sub>2</sub>CO, H<sub>2</sub>CS, c-HCOOH, t-HCOOH, and HNCO were chosen, most of which are well characterized experimentally. Furthermore, as far as we are concerned, some of the theoretical parameters of c-HCOOH and HNCO are being reported for the first time in this work.

In order to simulate rotational spectra, it is necessary at first to determine the spectroscopic parameters, which are subsequently employed in a simulation of rotational transitions by means of a model Hamiltonian matrix diagonalization. In the section "Theoretical background," some of the theoretical background are discussed with a particular emphasis on the spectroscopic parameters and the necessary approximations to the complete Hamiltonian which are required to simulate the rotational spectra. In the section "Computational details," the computational details are presented. In the section "Results and discussion," results are reported and discussed in detail. Finally, an overview and conclusions are presented in the section "Conclusions."

#### **Theoretical background**

#### The asymmetric rigid rotor

The rotational energy of a rigid rotor is classically expressed as:

$$T_{rot} = \frac{1}{2} (I_a \omega_a^2 + I_b \omega_b^2 + I_c \omega_c^2),$$
(1)

where  $\omega_a$ ,  $\omega_b$ , and  $\omega_c$  are the angular velocities of the system with respect to the main molecule-fixed axes *a*, *b*, and *c*. The rigid rotor Hamiltonian can therefore be described as a kinetic energy operator expressed in terms of the components of the angular momentum  $\hat{\mathbf{J}}$  with relation to the main axes [42]:

$$\hat{H}_{rot} = \frac{1}{2} \frac{\hat{J}_a^2}{I_a} + \frac{1}{2} \frac{\hat{J}_b^2}{I_b} + \frac{1}{2} \frac{\hat{J}_c^2}{I_c}$$
(2)

The energies of the rotational levels are the eigenvalues of the Hamiltonian in expression Eq. 2. The unknown asymmetric top eigenfunctions  $|\psi_i\rangle$  can be expressed as a linear expansion in terms of a known complete orthonormal set  $\{\varphi_i\}$  ( $\varphi_i \equiv |J', M', K'\rangle$ ), which is conveniently chosen as the symmetric top eigenfunctions:

$$|\psi_{i}\rangle = \sum_{J'} \sum_{M'} \sum_{K'} c_{J',M',K'}^{(i)} |J',M',K'\rangle$$
(3)

The sum can be simplified by considering only terms in which the symmetric top quantum numbers J' and M'have the same values of the asymmetric top eigenfunctions  $(|\psi_i\rangle)$ , namely J and M [41]:

$$|\psi_i\rangle = \sum_{K'=-J}^{J} c_{J,M,K'}^{(i)} |J, M, K'\rangle$$
 (4)

The eigenvalues of  $\hat{H}_{rot}$  (Eq. 2) are thus obtained by solving the resulting secular equation, where the symmetric top eigenfunctions  $\varphi_i$  were applied:

$$det[\langle \varphi_m | \hat{H}_{rot} | \varphi_n \rangle - E_i \delta_{mn}] = 0$$
<sup>(5)</sup>

The quantum number K' gives the angular momentum component along the molecule-fixed axis c for the symmetric top:

$$J_c|J, M, K'\rangle = \pm K'|J, M, K'\rangle \tag{6}$$

For asymmetric tops, the double degeneracy of the  $\pm K'$  symmetric top quantum numbers is removed, resulting in a splitting of the rotational energy levels and therefore much more complex rotational spectrum profiles [40].

## The rotational Hamiltonian beyond the rigid rotor approximation

For simulation purposes in an astrochemical context, corrections to the rotational Hamiltonian beyond the rigidrotor and harmonic-oscillator approximations ought to be accounted for. The effective rotational Hamiltonian for a given vibrational state is derived from the perturbative treatment of the complete vibrational-rotational Hamiltonian as described by Watson [44], resulting in a transformed Hamiltonian averaged over all vibrational coordinates [45]. This procedure leads to a separation of the rotational and vibrational motions [46], analogous to the splitting between the electronic and vibrational terms of the energy expression in the Born-Oppenheimer approximation.

Hence, the rotational Hamiltonian becomes a function of only the rotational angular momentum operators and vibrationally dependent rotational constants. This effective Hamiltonian can be expressed as a power law whose coefficients are the rotational and centrifugal distortion constants at a particular vibrational state [47–49]:

$$\frac{H_{rot}}{hc} = \sum_{\alpha} B_{\nu}^{(\alpha)} J_{\alpha}^{2} + 1/4 \sum_{\alpha,\beta} (\tau_{\alpha\alpha\beta\beta}^{'})_{\nu} J_{\alpha}^{2} J_{\beta}^{2} + \sum_{\alpha} \Phi_{\alpha\alpha\alpha} J_{\alpha}^{6} \\
+ \sum_{\alpha\neq\beta} \Phi_{\alpha\alpha\beta} (J_{\alpha}^{4} J_{\beta}^{2} + J_{\beta}^{2} J_{\alpha}^{4}) + \dots$$
(7)

where  $\alpha$  and  $\beta$  denote the three main molecule-fixed axes, namely *a*, *b*, and *c*. Ergo,  $J_{\alpha}$  are the components of the total angular momentum along the three main axes in units of  $\hbar$ . The operator hat notation is henceforth omitted for the sake of simplicity.  $B_{\nu}^{(\alpha)}$  corresponds to the effective rotational constants,  $(\tau'_{\alpha\alpha\beta\beta})_{\nu}$  to the effective quartic centrifugal distortion constants as defined by Kivelson and Wilson [50] and  $\Phi_{\alpha,\beta,\gamma}$  to the effective sextic centrifugal distortion constants as defined by Watson [48].

The effective Hamiltonian (Eq. 7) can be further simplified by means of contact transformations, since the eigenvalues of a Hamiltonian are unchanged by those operations [51]. Hence, a reduced Hamiltonian [48, 51–53], containing only the linear combinations of the coefficients in Eq. 7, and indistinguishable from the latter in terms of energy levels, is derived.

The reduced Hamiltonian for asymmetric tops as derived by Watson is [51]:

$$\frac{H_{rot}^{A}}{hc} = \sum_{\alpha} B_{\nu}^{(\alpha)} J_{\alpha}^{2} - \Delta_{J} J^{4} - \Delta_{JK} J^{2} J_{z}^{2} - \Delta_{K} J_{z}^{4} - 2\delta_{J} J^{2} (J_{x}^{2} - J_{y}^{2}) 
-\delta_{K} [J_{z}^{2} (J_{x}^{2} - J_{y}^{2}) + (J_{x}^{2} - J_{y}^{2}) J_{z}^{2}] + H_{J} J^{6} + H_{JK} J^{4} J_{z}^{2} 
+ H_{KJ} J^{2} J_{z}^{4} + H_{K} J_{z}^{6} + 2h_{J} J^{4} (J_{x}^{2} - J_{y}^{2}) + h_{JK} J^{2} [J_{z}^{2} (J_{x}^{2} - J_{y}^{2}) 
+ (J_{x}^{2} - J_{y}^{2}) J_{z}^{2}] + h_{K} [J_{z}^{4} (J_{x}^{2} - J_{y}^{2}) + (J_{x}^{2} - J_{y}^{2}) J_{z}^{4}] + \dots \qquad (8)$$

where  $B_{\nu}^{(a)} \equiv A_{\nu}$ ,  $B_{\nu}^{(b)} \equiv B_{\nu}$ , and  $B_{\nu}^{(c)} \equiv C_{\nu}$ . This reduced Hamiltonian is commonly addressed as Watson's A-type Hamiltonian. For the special case where the rotational constants are accidentally quasi-degenerate and the system approaches a symmetric top, Watson's A Hamiltonian fails [51]. It is thus necessary to employ a different form of the reduced Hamiltonian, namely Watson's S Hamiltonian [49]:

$$\frac{\tilde{H}_{rot}^{S}}{hc} = \sum_{\alpha} B_{v}^{(\alpha)} J_{\alpha}^{2} - D_{J} J^{4} - D_{JK} J^{2} J_{z}^{2} - D_{K} J_{z}^{4} + d_{1} J^{2} (J_{+}^{2} + J_{-}^{2}) + d_{2} (J_{+}^{4} + J_{-}^{4}) + H_{J} J^{6} + H_{JK} J^{4} J_{z}^{2} + H_{KJ} J^{2} J_{z}^{4} + H_{K} J_{z}^{6} + h_{1} J^{4} (J_{+}^{2} + J_{-}^{2}) + h_{2} J^{2} (J_{+}^{4} + J_{-}^{4}) + h_{3} (J_{+}^{6} + J_{-}^{6}) + \dots (9)$$

#### Vibration-rotation interaction constants

For simplicity, some equations in this section will be expressed solely for the rotational constant B along the b axis, albeit analogous for the other main axes. The asymmetric top rotational constant  $B_{\nu}$  corrected for the vibrational-rotational coupling of the molecule is expressed as [45]:

$$B_{\nu} = B_e - \sum_r \alpha_r^B \left(\nu_r + \frac{1}{2}\right) + \dots \tag{10}$$

where  $B_e$  denotes the equilibrium rotational constant along the b axis:

$$B_e = \hbar^2 / 2hcI_b \tag{11}$$

The sum in Eq. 10 includes all vibrational normal modes of the system. The vibrational-rotational coupling constant along the b axis  $\alpha_r^B$  is derived as [45]:

$$-\alpha_{r}^{B} = \frac{2B_{e}^{2}}{\omega_{r}} \left[ \sum_{\xi} \frac{3(a_{r}^{(b\xi)})^{2}}{4I_{\xi}} + \sum_{s \neq r} (\zeta_{r,s}^{(b)})^{2} \frac{(3\omega_{r}^{2} + \omega_{s}^{2})}{\omega_{r}^{2} - \omega_{s}^{2}} + \pi \left(\frac{c}{h}\right)^{1/2} \sum_{s} \phi_{rrs} a_{s}^{(bb)} \left(\frac{\omega_{r}}{\omega_{s}^{3/2}}\right) \right]$$
(12)

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where  $\omega_r$  is the  $r^{th}$  harmonic vibrational frequency, corresponding to the  $Q_r$  normal mode. The inertial derivatives  $a_r^{(\alpha\beta)}$  are defined as:

$$a_r^{(\alpha\beta)} = (\partial I_{\alpha\beta}/\partial Q_r)_e \tag{13}$$

where  $I_{\alpha\beta}$  denotes equilibrium moments and products of inertia [54]. The Coriolis zeta constant  $\zeta_{r,s}^{(\alpha)}$  couples the normal modes  $Q_r$  and  $Q_s$  through rotation about the  $\alpha$  axis:

$$\zeta_{r,s}^{(\alpha)} = \sum_{i} \left( L_{ir}^{(\beta)} L_{is}^{(\gamma)} - L_{ir}^{(\gamma)} L_{is}^{(\beta)} \right)$$
(14)

where **L** corresponds to a matrix that transforms the normal coordinates to mass-weighted Cartesian coordinates. The  $\phi_{rst}$  constants in the third sum of Eq. 12 denote the cubic anharmonic force constants derived from a Taylor expansion of the vibrational potential in terms of the dimensionless normal coordinates  $q_r$ , as described by Nielsen [55–57]:

$$\frac{V}{hc} = \frac{1}{2} \sum_{r} \omega_r q_r^2 + \frac{1}{6} \sum_{rst} \phi_{rst} q_r q_s q_t + \dots$$
(15)

Thus, the vibrational-rotational coupling can be split between two different contributions, the  $a_r^{(\alpha\beta)}$  and  $\zeta_{r,s}^{(\alpha)}$ constants, which are combined together in a resulting vibrational-rotational parameter  $a_r^{(\alpha)}$  (Eq. 12). For an extensive number of molecules, many of the cubic force constants related to stretching modes have negative values, which dominate the expression of  $a_r^{(\alpha)}$  for small systems. For this reason, the left side of Eq. 12 is conveniently defined as negative. We should also note that, when two vibrational states are accidentally quasi-degenerate, the Coriolis term in Eq. 12 breaks down, since its perturbative treatment is no longer valid. This configures a special case known as Coriolis resonance, which can be partially solved by neglecting the first-order terms and instead replacing them by [57–59]:

$$(\zeta_{r,s}^{(b)})^2 \frac{(3\omega_r^2 + \omega_s^2)}{\omega_r^2 - \omega_s^2} \to -\frac{1}{2} (\zeta_{r,s}^{(b)})^2 \frac{(\omega_r - \omega_s)^2}{\omega_s(\omega_r + \omega_s)}$$
(16)

#### Anharmonic vibration constants

The energy levels of a polyatomic molecule can be empirically described by the sum of a vibrational term independent of the rotational quantum numbers and a rotational term parametrically dependent on the vibrational quantum numbers [45]:

$$T(v, J) = G(v) + F_v(J)$$
 (17)

Similar to the procedure described in the section "The rotational Hamiltonian beyond the rigid rotor approximation," G(v) can be derived by means of a perturbative treatment of the expanded vibrational potential in Eq. (16):

$$G(v) = \sum_{r} \omega_r \left( v_r + \frac{1}{2} \right) + \sum_{r} \sum_{s \le r} \chi_{rs} \left( v_r + \frac{1}{2} \right) \left( v_s + \frac{1}{2} \right) + \dots$$
(18)

where  $\chi_{rs}$  are the anharmonic vibrational constants derived by Nielsen [57]. Following the notation in [45], the diagonal terms of the  $\chi_{rs}$  matrix are:

$$\chi_{rr} = \frac{1}{16}\phi_{rrrr} - \frac{1}{16}\sum_{s}\phi_{rrs}^{2} \left[\frac{8\omega_{r}^{2} - 3\omega_{s}^{2}}{\omega_{s}(4\omega_{r}^{2} - \omega_{s}^{2})}\right]$$
(19)

And for  $r \neq s$ :

$$\chi_{rs} = \frac{1}{4} \phi_{rrss} - \frac{1}{4} \sum_{t} \left( \frac{\phi_{rrt} \phi_{tss}}{\omega_{t}} \right) - \frac{1}{2} \sum_{t} \left[ \frac{\phi_{rst}^{2} \omega_{t} (\omega_{t}^{2} - \omega_{r}^{2} - \omega_{s}^{2})}{\Delta_{rst}} \right] + \left[ A_{e} (\zeta_{r,s}^{(a)})^{2} + B_{e} (\zeta_{r,s}^{(b)})^{2} + C_{e} (\zeta_{r,s}^{(c)})^{2} \right] \left[ \frac{\omega_{r}}{\omega_{s}} + \frac{\omega_{s}}{\omega_{r}} \right] (20)$$

where:

$$\Delta_{rst} = (\omega_r + \omega_s + \omega_t)(\omega_r - \omega_s - \omega_t)(\omega_r + \omega_s - \omega_t)(\omega_r - \omega_s + \omega_t)$$
(21)

Those constants can be employed to derive frequencies with the best correspondence to the experimental values, known as the fundamental frequencies  $v_r$ . For an asymmetric top,  $v_r$  are given by the expression [58]:

$$v_r = \omega_r + \Delta_r = \omega_r + 2\chi_{rr} + \frac{1}{2}\sum_{s \neq r} \chi_{rs} + \dots$$
 (22)

#### **Centrifugal distortion constants**

Wilson and Howard [46] described the centrifugal distortion effects as the following term in the rotational Hamiltonian resulting from the perturbative treatment:

$$\sum_{\alpha\beta\gamma\delta} \frac{1}{4} \tau_{\alpha\beta\gamma\delta} J_{\alpha} J_{\beta} J_{\gamma} J_{\delta}$$
(23)

where  $\tau_{\alpha\beta\gamma\delta}$  corresponds to:

$$\tau_{\alpha\beta\gamma\delta} = -\frac{1}{2} \sum_{r} \left( \frac{a_r^{(\alpha\beta)} a_r^{(\gamma\delta)}}{\lambda_r I_\alpha I_\beta I_\gamma I_\delta} \right)$$
(24)

Notably, the centrifugal distortion constants (Eq. 24) are, to a good approximation, only functions of the quadratic vibrational potential [52, 58]. The terms in Eq. 24 that have a first-order contribution to the rotational Hamiltonian can be expressed as the centrifugal distortion constant  $\tau'_{\alpha\alpha\beta\beta}$  [51, 60, 61]:

$$\tau'_{\alpha\alpha\beta\beta} = \tau_{\alpha\alpha\beta\beta} + 2\tau_{\alpha\beta\alpha\beta}(1 - \delta_{\alpha\beta}) \tag{25}$$

Watson demonstrated that only five linear combinations of the six  $\tau_{\alpha\beta\gamma\delta}$  constants are directly measurable from experimental spectrum [51, 61]. Those linear combinations are expressed in Watson's A Hamiltonian (Eq. 8) as the five independent quartic centrifugal distortion constants:

$$\Delta_J = D_J - 2R_6$$
  

$$\Delta_{JK} = D_{JK} + 12R_6$$
  

$$\Delta_K = D_K - 10R_6$$
  

$$\delta_J = \delta_J$$
  

$$\delta_K = -2R_5 - 4\sigma R_6$$
(26)

The parameters in Eq. 26 are derived by Kivelson and Wilson [60], and are discussed in more detail in Supplementary Information (SI).

In order to obtain the sextic centrifugal distortion correction, it is necessary to perform a perturbative treatment to the fourth order of the complete vibrational-rotational Hamiltonian [44, 62–65]. The ten resulting centrifugal distortion constants can be then rearranged into seven observable independent linear combinations [48, 52, 66]:

$$H_J = \Phi_{600} + 2\Phi_{204} \tag{27a}$$

$$H_{JK} = \Phi_{420} - 12\Phi_{204} + 2\Phi_{024} + 16\sigma\Phi_{006} - 16(R_5 - 2\sigma R_6)s_{111} + 8(2A'_0 - B'_0 - C'_0)s_{111}^2$$
(27b)

$$H_{KJ} = \Phi_{240} + \frac{10}{3}\Phi_{420} - 30\Phi_{204} - \frac{10}{3}H_{JK}$$
(27c)

$$H_K = \Phi_{060} - \frac{7}{3}\Phi_{420} + 28\Phi_{204} + \frac{7}{3}H_{JK}$$
(27d)

$$h_J = \Phi_{402} + \Phi_{006}$$
(27e)  
$$h_{IK} = \Phi_{222} - 10\Phi_{006} + 4\sigma\Phi_{204} + 2(D_{IK} - 2\sigma\delta_I - 4R_6)s_{111}$$

$$h_{K} = \Phi_{042} + \frac{4}{3}\sigma \Phi_{024} + \left(9 + \frac{32}{3}\sigma^{2}\right) \Phi_{006} + 4\left[D_{K} - \frac{2}{3}\sigma R_{5} + 2\left(1 + \frac{8}{3}\sigma^{2}\right)R_{6}\right]s_{111} + \left(6 + \frac{10}{3}\sigma^{2}\right)\left(B'_{0} - C'_{0}\right)s_{111}^{2}$$
(27g)

The experimental spectra can be adjusted to a particular reduced Hamiltonian by means of the  $s_{111}$  parameter. In Watson's A Hamiltonian,  $s_{111}$  corresponds to:

$$s_{111} = -\frac{4R_6}{(B'_0 - C'_0)} \tag{28}$$

A detailed description of the  $\Phi_{2l,2m,2n}$  parameters can be found in tables IV and VII of the original work by Aliev and Watson [52]. Unlike for quartic centrifugal distortion constants, the sextic constants indeed depend on the cubic vibrational potential. Moreover, the contribution of centrifugal distortion to the effective rotational constants is briefly discussed in the SI.

#### **Transition intensities**

Rotational transitions are induced by the interaction of the species' permanent electric dipole moment with the radiation field. The intensities of those transitions are proportional to the square of the dipole moment matrix [67], precluding the observation of pure rotational lines from molecules with no permanent dipole moment. For asymmetric rotors, the dipole moment can be pointed to any arbitrary direction with relation to the main inertial axes. Consequently, the z components of the dipole moment matrix elements can be described as [40]:

$$\mu_{z} = \mu_{a} \langle \psi_{i} | cos(az) | \psi_{j} \rangle + \mu_{b} \langle \psi_{i} | cos(bz) | \psi_{j} \rangle + \mu_{c} \langle \psi_{i} | cos(cz) | \psi_{j} \rangle$$
(29)

Where  $\mu_a$ ,  $\mu_b$ , and  $\mu_c$  are the components of the electric dipole moment along the inertial axes, and  $\cos(az)$ ,  $\cos(bz)$ , and  $\cos(cz)$  correspond to the cosines of the angles between the main inertial axes and the space-fixed *z* axis. The wave functions for asymmetric rotors are denoted as  $\psi_i$ .

#### **Computational details**

In order to explore the influence of different functionals on rotational spectral simulations, geometry optimization and spectroscopic parameter calculations were performed with the Generalized Gradient Approximation (GGA) functional B97D [68, 69] and the Double-Hybrids (DH) functionals B2PLYP [70] and mPW2PLYP [71]. Grimme's D2 dispersion correction [69] was included in all functionals, and Dunning's aug-cc-pVTZ augmented basis set [72] was used throughout. Geometry optimizations were carried out with Berny's algorithm [73], with  $1.5 \times 10^{-5}$ Hartree/Bohr and  $6.0 \times 10^{-5}$  Angstrom convergence criteria (Gaussian keyword "Opt=Tight"). For all systems, an ultrafine (99,590) integration grid was used (Gaussian keyword "Int=UltraFine"). All calculations were performed with Gaussian 2009 quantum-chemical package [74] and PGO-PHER general purpose software [75] was employed on the rotational spectral simulations. All simulated spectra were obtained by considering rotational-vibrational coupling, at VPT2 level, anharmonic and quartic centrifugal-distortion corrections.

For comparison and validation purposes, experimental spectra of H<sub>2</sub>CO, H<sub>2</sub>CS, c-HCOOH, t-HCOOH, and HNCO were obtained from NASA's Jet Propulsion Laboratory

Table 1	Calculated	and experime	ental spectral	parameters of	of H <sub>2</sub> CO
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	B2PLYP	B2PLYPD	B97D	mPW2PLYPD	Experimental	Martin and Lee (1993)
	cc-pV1Z	aug-cc-pV1Z	aug-cc-pV1Z	aug-cc-pV1Z		
A <sub>0</sub> (MHz)	283,736.102	282,432.661	277,275.761	283,300.659	281,970.370 <sup>a</sup>	282,633.987
B <sub>0</sub>	38,766.147	38,714.414	38,526.953	38,912.345	38,835.426	38,465.171
C <sub>0</sub>	33,972.283	33,912.840	33,689.527	34,078.056	34,005.730	33,726.651
μ (D)	2.290	2.406	2.257	2.429	2.330 <sup>b</sup>	
$\Delta_J$ (MHz)	0.0731	0.0738	0.0746	0.0732	0.0755 <sup>a</sup>	0.0727
$\Delta_{JK}$	1.2778	1.2947	1.3257	1.2927	1.2936	1.2652
$\Delta_K$	18.8583	18.6700	19.0349	18.6516	19.9782	18.6093
$\delta_J$	0.0097	0.0098	0.0101	0.0097	0.0104	0.00956
$\delta_K$	0.9090	0.9179	0.9439	0.9162	1.0349	0.8963
$H_K \times 10^{-3}$ (MHz)	3.810	3.828	4.149	3.793		3.789
$H_{KJ} \times 10^{-5}$	1.112	1.440	0.7980	1.489		-5.810
$H_{JK} \times 10^{-5}$	2.554	2.605	2.794	2.588		2.580
$H_J \times 10^{-8}$	8.624	8.592	9.752	8.911		8.994
$h_J \times 10^{-8}$	4.042	4.102	4.460	4.126		2.998
$h_{JK} \times 10^{-5}$	1.368	1.395	1.492	1.388		1.370
$h_K \times 10^{-3}$	1.194	1.208	1.275	1.199		1.156

Theoretical values at CCSD(T) level of theory are taken from Martin and Lee (1993)[81]

<sup>a</sup>Experimental effective rotational constants and quartic centrifugal distortion constants obtained from [82] and values therein

<sup>b</sup>Experimental electric dipole moment from [83]

(JPL) database<sup>1</sup> and the Cologne Database for Molecular Spectroscopy (CDMS)<sup>2</sup>.

#### **Results and discussion**

#### H<sub>2</sub>CO

Formaldehyde (H<sub>2</sub>CO) was one of the first molecules to be detected with radiotelescopes [66], and has since been widely observed toward interstellar and circumstellar environments [76–79]. It is commonly used to study the physical conditions of astrophysical sources, acting as a probe for gas temperature and density [80]. In Table 1, calculated and experimental spectroscopic parameters of H<sub>2</sub>CO are presented. Theoretical rotational constants corrected for vibrational-rotational effects (hereafter "effective rotational constants," see Eqs. 10 and 12) show excellent agreement with experimental values, with mean discrepancies of 0.3% for all hybrid functionals and  $\sim 1\%$  for B97D. Centrifugal corrections to the rotational constants as described in equation S4 result in relative changes up to 0.001% for all case studies, and therefore were neglected in order to allow comparison with other theoretical works where only vibrational-rotational corrections were reported. For the quartic centrifugal distortion constants, mean errors of  $\sim$  4% were achieved. Moreover, both hybrid functionals yielded spectroscopic constants generally closer to experiment than CCSD(T) calculations from ref. [81]. Those results demonstrate that the methods and models used here describe very well the formaldehyde parameters and spectrum, with a slightly better performance of the B2PLYP functional with dispersion corrections. Although an already small mean error of  $\sim$  0.3% is achieved for the effective rotational constants at the B2PLYP/cc-pVTZ level of theory, with the addition of dispersion corrections and diffuse functions, this error is further improved to  $\sim 0.25\%$ . Similarly to B2PLYP, calculations with the mPW2PLYP functional with dispersion corrections yielded errors of  $\sim 0.29\%$ , evincing the generally good performance of hybrid functionals at spectroscopic parameter determinations. Overall, D2 dispersion corrections and augmented basis sets have shown better accuracy with respect to the measured constants, suggesting that those effects should be included in the theory even for tightly bounded molecules such as  $H_2CO$  [58]. Nevertheless, the B97D/aug-cc-pVTZ level yielded the largest errors for the effective rotational constants, and centrifugal distortion constants only slightly better than B2PLYP/cc-pVTZ, despite accounting for dispersion effects.

<sup>&</sup>lt;sup>1</sup>Available in https://spec.jpl.nasa.gov/ftp/pub/catalog/catdir.html

<sup>&</sup>lt;sup>2</sup>Available in https://cdms.astro.uni-koeln.de/cdms/portal/

Although all centrifugal distortion constants are generally difficult to derive theoretically, the  $\delta_K$  constant is specially sensitive to the level of theory [58], leading to systematically larger discrepancies in comparison with the other quartic constants. Moreover, as discussed by Clabo et al. [58], the experimental value for  $\delta_K$  is not completely consistent with the other measured constants. From the experimental values for  $\Delta_J$ ,  $\Delta_{JK}$ ,  $\Delta_K$ , and  $\delta_J$ , and by assuming planarity conditions [84], one can derive the experimental  $\tau$  constants that, subsequently, yield  $\delta_K$  = 0.9277 MHz. This value is much closer to the predicted  $\delta_K$  for all levels of theory, with a mean relative error smaller than 1%. Thus, the current theoretical formulation for the  $\delta_K$  constant indeed seems not to describe, to a good extent, the phenomenological parameter  $\delta_K$  obtained from experiments.

In what concerns spectral simulations, even slight differences in accuracy for the effective rotational constants are of great relevance to accurately predict line frequencies, since those constants have a zeroth-order contribution to the rotational energy. The simulated spectrum of formaldehyde at 300K with mPW2PLYPD/aug-cc-pVTZ level of theory is shown in Fig. 1, in comparison with the experimental spectrum from the JPL database. The simulation was truncated at a maximum upper J value of 57, and Watson's A Hamiltonian was employed. Sextic centrifugal distortion constants are difficult to determine both theoretically and experimentally, and thus were not included in the simulations. For transitions with lower Jvalues (< 800 GHz), discrepancies between experimental and theoretical frequencies with mPW2PLYPD functional were around  $\sim$ 1.00 GHz ( $\sim$  0.13% in relative terms). It is worth emphasizing that this frequency range corresponds to the relevant region for astronomical radioobservations. Comparatively, B2PLYP/cc-pVTZ (Fig. S1) and B2PLYPD/aug-cc-pVTZ (Fig. S2) yielded errors of 0.53 and  $\sim$  1.3 GHz, respectively. Although the  $\sim$ B2PLYP functional predicted the most accurate rotational frequencies for this system, the overall best performing functional for asymmetric tops was mPW2PLYP, as will be discussed in Sections H<sub>2</sub>CS, HCOOH, and HNCO. This apparently contradictory result can be understood by analyzing each effective rotational constant separately. For B2PLYPD/aug-cc-pVTZ, the A0 constant yielded a relative error of  $\sim 0.16\%$ , which was considerably closer to experiment than for B2PLYP/cc-pVTZ ( $\sim 0.62\%$ ) and mPW2PLYPD/aug-cc-pVTZ (~ 0.47%). However, the B<sub>0</sub> and C<sub>0</sub> constants predicted by B2PLYP with dispersion corrections were less accurate than both other levels of theory. In fact, the B2PLYP/cc-pVTZ description achieved the most accurate B<sub>0</sub> and C<sub>0</sub> constants, which consequently leads to better frequency predictions. Indeed, for formaldehyde, the A<sub>0</sub> effective rotational constant seems to be the least dominant regarding rotational spectrum simulations. This result demonstrates that one must be attentive of all rotational constants separately when performing spectrum simulations. It should also be noted that both  $B_0$  and  $C_0$  constants obtained from CCSD(T) calculations by ref. [81] were on average  $\sim 350$  MHz further from experiment in comparison with hybrid DFT results.

This outcome is suited to guide laboratory experiments of recording and analyzing microwave spectra which play a key role in the fields of radioastronomy and astrochemistry.

Fig. 1 Rotational spectrum of  $H_2CO$  at 300 K simulated with mPW2PLYPD/aug-cc-pVTZ level of theory (black). The experimental spectrum is shown in red. In the right panel, an astronomical relevant region is zoomed in



The simulated spectrum at B97D/aug-cc-pVTZ level of theory is shown in Fig. S3. The vibrational-rotational coupling constants (see Eq. 12) at the mPW2PLYP/aug-cc-pVTZ level of theory for all case studies are also listed in the SI.

#### H<sub>2</sub>CS

Since the 1970s, sulfur-containing molecules have been observed to be depleted in dense regions of the ISM in comparison with cosmic abundances [85-88], while in diffuse and highly ionized regions their abundances are seemingly cosmic [89–93]. Inasmuch as the sulfur chemical sink has yet to be identified, this phenomenon, namely the sulfur depletion problem [94], has been arousing much interest of the astrochemical community toward sulfur-bearing molecules. As a relevant example, thioformaldehyde  $(H_2CS)$  is a slightly asymmetric rotor that has been observed toward several interstellar sources [22, 95–99]. However similar in structure to its oxygen analogue H<sub>2</sub>CO, thioformaldehyde has much less optically thick lines, therefore being a good alternative to determine the physical properties of a particular source. Two main formation routes have been suggested for this molecule, resulting from either reactions involving atomic sulfur and

Table 2 Calculated and experimental spectral parameters of H<sub>2</sub>CS

the methyl radical (CH<sub>3</sub>) in hot cores [100], or forming on the surface of grain mantles within dense molecular clouds [88].

Theoretical and experimental spectroscopic parameters of H<sub>2</sub>CS are listed in Table 2. While all functionals derived effective rotational constants close to experiment (with relative mean errors less than 1%), B97D yielded slightly less accurate results than the hybrid functionals. However, its predictions of the quartic centrifugal distortion constants were the closest to experiment, again evincing an apparent advantage of the GGA functional in describing the quadratic force field of asymmetric tops. Martin et al. [101] have derived theoretical values for the spectroscopic constants of H<sub>2</sub>CS with CCSD(T) calculations, which are listed in Table 2. As for  $H_2CO$ , the parameters obtained through DFT calculations were comparable with the predictions of the coupled-cluster approach, but with much less demanding computational effort. The exceptions are the  $H_{KJ}$ ,  $h_{JK}$ , and  $h_K$  sextic centrifugal distortion constants, whose DFT predictions deviate more than one order of magnitude from the CCSD(T) calculations, possibly due to the higher difficulty associated with their derivation [102, 103].

From the experimental  $\Delta_J$ ,  $\Delta_{JK}$ ,  $\Delta_K$ , and  $\delta_J$ , we derive a new  $\delta_K = 0.3326$  MHz. For this new value, the mean

	B2PLYP cc-pVTZ	B2PLYPD aug-cc-pVTZ	B97D aug-cc-pVTZ	mPW2PLYPD aug-cc-pVTZ	Experimental	Martin et al. (1994)
$A_0$ (MHz)	293,888.436	293,079.382	289,947.613	294,311.184	291,613.519 <sup>a</sup>	293,547.182
$B_0$	17,593.668	17,574.772	17,476.466	17,656.413	17,699.7198	17,776.928
C <sub>0</sub>	16,563.637	16,544.104	16,446.043	16,620.665	16,651.7382	16,762.086
μ (D)	1.724	1.759	1.733	1.774	1.649 <sup>b</sup>	
$\Delta_J$ (MHz)	0.01867	0.01879	0.01917	0.01859	0.019358 <sup>a</sup>	0.0187
$\Delta_{JK}$	0.51451	0.51821	0.52730	0.51557	0.52095	0.51145
$\Delta_K$	22.37431	22.34474	22.94792	22.30371	23.3260	22.0590
$\delta_J$	0.00110	0.00111	0.00114	0.00110	0.00120457	0.00109
$\delta_K$	0.32309	0.32524	0.33313	0.32427	0.37195	0.31966
$H_K \times 10^{-3} (MHz)$	5.316	5.369	5.840	5.273	5.483	5.2311299
$H_{KJ} \times 10^{-5}$	-0.2482	-0.1005	-0.2700	-0.09278		-3.900519
$H_{JK} \times 10^{-6}$	3.744	3.783	4.036	3.756		3.56776
$H_J \times 10^{-10}$	11.05	9.543	8.529	15.52		2.5
$h_J \times 10^{-9}$	1.704	1.718	1.800	1.733		1.62
$h_{JK} \times 10^{-6}$	2.114	2.136	2.265	2.119		657.68837
$h_K \times 10^{-4}$	6.742	6.816	7.176	6.754		0.0201943

Theoretical values at CCSD(T) level of theory are taken from Martin et al. [101]

<sup>a</sup>Experimental effective rotational constants and quartic centrifugal distortion constants obtained from [104]

<sup>b</sup>Experimental electric dipole moment from [105]

relative errors of the  $\delta_K$  constants obtained with all three functionals corrected for dispersion effects ranged from 0.16 to 2.5%. Those discrepancies were systematically lessened by around 10% when compared with the measured constant, once more supporting the hypothesis of a faulty theoretical description of the  $\delta_K$  constant.

The best performing functional for H<sub>2</sub>CS was mPW2PLYP, which yielded mean errors of ~ 0.45% for the effective rotational constants and ~ 6.20% for the quartic centrifugal distortion constants. The simulated spectrum of thioformaldehyde at 300K with this functional is shown in Fig. 2, in comparison with the experimental spectrum from the JPL database. Watson's A Hamiltonian was employed in the simulation, with  $J_{max} = 27$ . Again, sextic centrifugal distortion constants were neglected. The simulated spectra at B2PLYPD/aug-cc-pVTZ and B97D/aug-cc-pVTZ levels of theory are shown in the SI.

For *J* values up to 10, simulated transitions at the mPW2PLYPD/aug-cc-pVTZ level of theory yielded errors ranging from ~ 200 to ~ 675 MHz. Comparatively, the mean discrepancies for the same low *J* interval at B2PLYPD and B97D levels of theory were, respectively, ~ 1500 MHz and ~ 2700 MHz. Notably, accurate predictions of the rotational constants are fundamental to the spectrum simulations. In particular, the B<sub>0</sub> and C<sub>0</sub> constants have repeatedly proven to play a key role on the frequency calculation of the rotational transitions of molecules with C<sub>2v</sub> symmetry. Moreover, the B<sub>0</sub> and C<sub>0</sub> rotational constants obtained with CCSD(T) calculations

were considerably further from experiment than those predicted by the mPW2PLYPD functional.

#### HCOOH

The first detections of formic acid (HCOOH) in the interstellar medium were reported by Zuckerman et al. (1971) [106] and Winnewisser and Churchwell (1975) [107] toward the Giant Molecular Cloud (GMC) Sgr B2. Thenceforth, both rotational isomers have been detected toward a variety of sources such as high- and low-mass star-forming regions and protoplanetary disks [108-110]. Its structural similarities to more complex, biologically important species such as amino acids give formic acid a distinctive role in the field of astrobiology. Indeed, chemical routes to form glycine from HCOOH have been explored since the 1970s (e.g., [111-113]). In the present work, we have obtained the spectroscopic parameters and simulated rotational spectra for the cis and trans rotamers of HCOOH. Both isomers have  $C_s$ -symmetry, with an energy difference of 16.3 kJ mol<sup>-1</sup> between them [114]. It should be noted, however, that since both rotamers are in equilibrium at 300K, their rotational spectrum measurements are virtually inseparable. The determination of their respective rotational frequencies and spectroscopic parameters from experiment must then be performed by means of separate fittings of the intensity-weighted lines within the conjoined spectrum. Thus, caution is necessary when comparing experiment and theory for these species.



Fig. 2 Rotational spectrum of H<sub>2</sub>CS at 300 K simulated with mPW2PLYPD/aug-cc-pVTZ level of theory (black). The experimental spectrum is shown in red. In the right panel, an astronomical relevant region is zoomed in

Table 3 presents the calculated and experimental spectroscopic parameters of c-HCOOH. Considering the main terms of the rotational energy expression, the best performance was achieved by the mP2WPLYPD/aug-ccpVTZ level of theory, with a relative error of  $\sim 0.29\%$ for the effective rotational constants and  $\sim$  2.9% for the quartic centrifugal distortion constants. Comparatively, the B2PLYP/cc-pVTZ and B2PLYPD/aug-cc-pVTZ levels of theory yielded errors of, respectively,  $\sim 0.54\%$ and  $\sim 0.60\%$  for the effective rotational constants. While the incorporation of dispersion corrections and diffuse functions to the B2PLYP functional resulted in a difference of  $\sim 460$  MHz in the rotational constant A<sub>0</sub> toward the experimental value, it also led to  $B_0$  and  $C_0$ values  $\sim 10$  MHz further from experiment. Conversely, mPW2PLYPD/aug-cc-pVTZ obtained both B<sub>0</sub> and C<sub>0</sub> constants  $\sim$  45 MHz closer than B2PLYP/cc-pVTZ to the measurements. Moreover, dispersion corrections have considerably improved the quartic centrifugal distortion constants with all functionals, as one would expect. The B97D/aug-cc-pVTZ level of theory yielded quartic centrifugal distortion constants closest to experiment, with a mean error of  $\sim 2.4\%$ . However, its larger discrepancies for the effective rotational constants ( $\sim 1.7\%$ ) make this functional the least adequate to perform spectrum simulations. Regardless of the choice of functional, however, the sextic centrifugal distortion constants show

Table 3 Calculated and experimental spectral parameters of c-HCOOH

great variance on accuracy (with relative errors ranging from  $\sim 11$  to  $\sim 100\%$  depending on the coefficient). Nonetheless, considering the small absolute magnitude of those constants and their high dependency on the theory level against a relatively small contribution to the rotational energy, higher errors are expected and rather tolerated.

Predicted and measured constants for the t-HCOOH are listed in Table 4. In this work, the effective rotational constants of the trans rotamer were also best predicted by the mPW2PLYPD/aug-cc-pVTZ level of theory, again with a mean relative error of  $\sim 0.29\%$ . Comparatively, the effective rotational constants obtained by Demaison et al. [115] with CCSD(T)/cc-pVTZ calculations yielded relative errors of  $\sim 0.7\%$ . As for the *cis* rotamer, B<sub>0</sub> and C<sub>0</sub> rotational constants calculated with mPW2PLYPD were displaced around 40 MHz toward the experimental value in comparison with B2PLYPD. For this configuration, the addition of diffuse functions and dispersion corrections to the B2PLYP functional were much less impactful than for the cis rotamer, only improving the A<sub>0</sub> rotational constant in the B2PLYPD/aug-cc-pVTZ level of theory by  $\sim 100$  MHz. Nevertheless, distinctions between the behaviors of each isomer could be largely influenced by the difficulty in dissociating their experimental spectra. The GGA functional yielded effective rotational constants with a mean relative error of  $\sim 1.8\%$ , the worst performance between all functionals. Regarding centrifugal effects, the

	B2PLYP cc-pVTZ	B2PLYPD aug-cc-pVTZ	B97D aug-cc-pVTZ	mPW2PLYPD aug-cc-pVTZ	Experimental <sup>a</sup>
A <sub>0</sub> (MHz)	86,703.563	86,242.515	85,059.319	86,768.958	86,461.624
$B_0$	11,604.990	11,596.418	11,472.794	11,654.416	11,689.185
C <sub>0</sub>	10,211.191	10,208.034	10,094.924	10,260.528	10,283.996
μ (D)	3.858	3.908	3.755	3.941	3.790
$\Delta_J$ (MHz)	0.00813	0.00824	0.00860	0.00816	0.00836
$\Delta_{JK}$	-0.06967	-0.06979	-0.06742	-0.07026	-0.07144
$\Delta_K$	2.38676	2.38092	2.39963	2.38177	2.36167
$\delta_J$	0.00136	0.00138	0.00143	0.00137	0.00142
$\delta_K$	0.03808	0.03841	0.04021	0.03813	0.04075
$H_K \times 10^{-4}$	1.639	1.644	1.722	1.629	1.851
$H_{KJ} \times 10^{-6}$	-4.348	-4.466	-4.764	-4.338	-9.673
$H_{JK} \times 10^{-7}$	-3.751	-3.891	-4.485	-3.787	-2.974
$H_J \times 10^{-9}$	5.768	5.719	5.377	5.798	10.64
$h_J \times 10^{-9}$	5.071	5.180	5.525	5.129	2.317
$h_{JK} \times 10^{-8}$	5.529	5.331	5.073	5.511	-73.00
$h_K \times 10^{-5}$	1.431	1.440	1.499	1.428	

Table 4 Calculated and experimental spectral parameters of t-HCOOH

	B2PLYP cc-pVTZ	B2PLYPD aug-cc-pVTZ	B97D aug-cc-pVTZ	mPW2PLYPD aug-cc-pVTZ	Experimental <sup>a</sup>	CCSD(T) <sup>b</sup> cc-pVTZ
A <sub>0</sub> (MHz)	77,194.408	77,021.950	76,693.034	77,300.596	77,512.235	77,377.000
$B_0$	11,978.294	11,953.367	11,775.965	12,023.530	12,055.106	11,934.067
C <sub>0</sub>	10,345.289	10,323.635	10,184.043	10,381.245	10,416.115	10,322.786
μ (D)	1.465	1.511	1.506	1.528	1.420	
$\Delta_J$ (MHz)	0.00836	0.00834	0.00838	0.00832	0.009996	0.00973
$\Delta_{JK}$	0.15283	0.15321	0.17484	0.15033	-0.086249	-0.08872
$\Delta_K$	1.2160	1.2321	1.3088	1.2212	1.702447	1.69560
$\delta_J$	0.00117	0.00116	0.00110	0.00117	0.001949	0.00189
$\delta_K$	0.09748	0.09723	0.10243	0.09654	0.042732	0.03863
$H_K \times 10^{-5}$	4.589	4.682	5.037	4.626	12.12	11.73
$H_{KJ} \times 10^{-6}$	1.373	1.066	1.686	1.089	-10.57	-9.2
$H_{JK} \times 10^{-6}$	1.212	1.241	1.367	1.217	0.102	0.013
$H_J \times 10^{-9}$	-2.394	-1.975	-2.838	-1.872	13.14	10.3
$h_J \times 10^{-9}$	1.166	1.349	1.247	1.360	5.763	7.3
$h_{JK} \times 10^{-7}$	6.115	6.379	7.162	6.236	0.9750	1.3
$\frac{h_K \times 10^{-5}}{2}$	2.567	2.575	2.841	2.531	1.482	1.3

### <sup>a</sup>[116]

<sup>b</sup>[115]

B2PYLP/aug-cc-pVTZ level of theory predicted the most accurate quartic centrifugal distortion constants, although for the t-HCOOH rotamer the calculated values were much farther from the experimental data than for c-HCOOH. In particular, the  $\Delta_{JK}$  and  $\delta_K$  constants predictions yielded relative errors of more than 100%, whereas for the other quartic constants the relative errors were around 25%. Furthermore, Demaison et al. [115] obtained quartic and sextic centrifugal distortion constants comparatively closer to experiments.

As for H<sub>2</sub>CO, the inconsistency of the  $\delta_K$  constant was investigated. From the planarity conditions and the experimental values for the other quartic centrifugal distortion constants, one can obtain  $\delta_K = 0.0378$  MHz for c-HCOOH and  $\delta_K = 0.0400$  MHz for t-HCOOH. For the *cis* rotamer, the derived experimental value of  $\delta_K$  was again generally closer to the theoretical predictions, yielding mean relative errors of  $\sim 1\%$ . The exception was the B97D/augcc-pVTZ level of theory, which predicted a  $\delta_K$  constant that was closer to measurements. For the trans isomer, however, the  $\tau$  constants (Eq. 24) seem to be less accurately described by the theory, and as a result the discrepancies of the  $\Delta_{JK}$  and  $\delta_K$  predictions were unusually large. Thus, the "corrected" value for  $\delta_K$  is still rather different from the calculations, since the latter are directly affected by the seemingly amiss  $\tau$  constants.

The simulated spectra of c-HCOOH and t-HCOOH at the mPW2PLYP/aug-cc-pVTZ level of theory are shown in Figs. 3 and 4, while the spectra simulated at the B2PLYPD/aug-cc-pVTZ and B97D/aug-cc-pVTZ levels of theory are presented in the SI. For the c-HCOOH rotamer, experimental transitions were obtained from the CDMS catalog, while for t-HCOOH the experimental spectrum was available at JPL. All simulations were performed with Watson's A-type Hamiltonian, truncating a J maximum of 100 for the *cis* and 20 for the *trans* rotamer.

For c-HCOOH at low J values (< J = 10), predicted frequency errors were around 100 MHz for the mPW2PLYP/aug-cc-pVTZ level of theory, ranging from  $\sim 13$  to  $\sim 195$  MHz. The mean errors associated with B2PLYP/cc-pVTZ and B2PLYPD/aug-cc-pVTZ were, respectively,  $\sim$  348 and  $\sim$  971 MHz. These results suggest that, similarly to H<sub>2</sub>CO and H<sub>2</sub>CS, the B<sub>0</sub> and C<sub>0</sub> effective rotational constants are much more relevant to the rotational energy levels of c-HCOOH than the A<sub>0</sub> constant. Although the B97D functional was the best for describing the quartic centrifugal distortion constants, frequency errors for this system even at low J values were approximately 3 GHz, much larger than for all the other levels of theory. Ergo, the larger discrepancies in the effective rotational constants predicted with B97D have much more influence than centrifugal corrections to the energy levels. Again,

**Fig. 3** Rotational spectrum of c-HCOOH at 300 K simulated with mPW2PLYPD/aug-ccpVTZ level of theory (black). In the right panel, an astronomical relevant region is zoomed in. The experimental spectrum is shown in red



it becomes evident that accurate rotational constants are fundamental to correctly predict rotational frequencies.

For t-HCOOH, the mPW2PLYP/aug-cc-pVTZ level of theory predicted frequencies with a mean error of 614 MHz at low J values, and B2PLYPD/aug-cc-pVTZ yielded errors of  $\sim$  1.47 GHz. Notably, frequency discrepancies for t-HCOOH were significantly higher than for the *cis* rotamer, despite both configurations presenting effective rotational constants similarly well described by theory. Thus, it is likely that the substantial difference between frequency prediction errors for each rotamer is due to the effectiveness of the functionals to describe the system's quadratic force field, and consequently to determine its centrifugal distortion effects. Nevertheless, the mPW2PLYP/aug-cc-pVTZ has consistently achieved the best theoretical frequencies for HCOOH, with sufficiently good predictions to guide microwave spectroscopy measurements.

For both HCOOH isomers, but especially for the *cis* rotamer, transition intensity predictions deviated considerably from the measured spectra. This result is unexpected, since both yielded permanent dipole moments fairly close to experiments with all functionals. Precisely, while the experimental values for the two dipole moment components of c-HCOOH are listed as  $\mu_a = 2.65$  D and



 $\mu_b = 2.71$  D [114], the mPW2PLYPD/aug-cc-pVTZ calculations derived  $\mu_a = 2.80$  D and  $\mu_b = 2.77$  D, the least accurate among all functionals. Indeed, semi-empirical simulations of c-HCOOH rotational spectrum where the experimental dipole moment components were employed also result in the same spectral profile as derived with the theoretical parameters.

#### HNCO

Isocyanic acid (HNCO), the smallest molecule containing the four biogenic elements, was first detected in the interstellar medium toward the GMC Sgr B2 [117–119]. Since then, it has been detected toward multiple different environments, ranging from high- and low-mass protostars to extragalactic sources [22, 120–124]. The main formation pathway for this molecule in hot cores consists in gasphase neutral-neutral reactions: complex organic molecules formed from HNCO evaporate from grain mantles to the gas phase, where they suffer dissociation processes leading to the production of HNCO [122, 125]. As a consequence, numerous works have suggested that HNCO could trace large-scale shocks, giving insight on the source's morphology [125–127].

In Table 5, calculated and experimental spectroscopic constants of HNCO are listed. The mean relative errors of

Table 5 Calculated and experimental spectral parameters of HNCO

the effective rotational constants yielded by the functionals B2PLYPD, B97D and mPW2PLYPD were, respectively,  $\sim 4.2\%$ ,  $\sim 6.1\%$ , and  $\sim 3.1\%$ . Those unusually large errors are predominantly associated with the A<sub>0</sub> constant, which was generally poorly predicted in comparison with the other case studies. As for H<sub>2</sub>CS and HCOOH, the best predictions were achieved by the mPW2PLYPD/aug-cc-pVTZ level of theory, whereas the effective rotational constants calculated with the GGA functional were the furthest from experiment. Nonetheless, all three functionals yielded better predictions than calculations at the CISD level of theory reported by Ref. [128]. Centrifugal distortion is large for HNCO, and as a consequence its constants were particularly difficult to predict, with discrepancies as big as three orders of magnitude compared with the measurements.

Isocyanic acid constitutes a quasi-symmetric rotor, given that the B<sub>0</sub> and C<sub>0</sub> rotational constants are accidentally quasi-degenerate. Thus, the S-reduction of Watson's Hamiltonian [49] was employed in the simulation. Figure 5 shows the predicted spectrum of HNCO at the mPW2PLYP/augcc-pVTZ level of theory and the experimental spectrum from the JPL database. The simulation was performed at T= 300K, with  $J_{max} = 47$  and sextic centrifugal distortion constants were not included. Intensity units were set to arbitrary due to an overflow of the Boltzmann expression at high J levels, which is a direct consequence of the large

	DADI VD		<b>D</b> 07D		Europeine entel	CIED
	cc-pVTZ	aug-cc-pVTZ	aug-cc-pVTZ	aug-cc-pVTZ	Experimental	CISD
A <sub>0</sub> (MHz)	1,006,272.618	1,022,565.218	1,075,036.194	994,939.517	918,504.400 <sup>a</sup>	818,424.000
B <sub>0</sub>	11,145.126	11,138.053	11,089.834	11,179.388	11,071.008	10,911.000
C <sub>0</sub>	10,861.358	10,851.251	10,785.718	10,904.344	10,910.576	10,768.000
μ (D)	2.109	2.079	2.034	2.090	2.070 <sup>b</sup>	
$\Delta_J$ (MHz)	0.00260	0.00259	0.00251	0.00265	0.003486 <sup>a</sup>	
$\Delta_{JK}$	15.9048	16.4925	18.3726	15.1341	0.931700	
$\Delta_K$	5906.15	6428.21	8876.09	5034.28	6065.60	
$\delta_J$	0.00027	0.00028	0.00034	0.00024	-0.000073	
$\delta_K$	0.00009	-0.00009	-0.00010	-0.00008	-0.000037	
$H_K \times 10^{+2} (MHz)$	1.243	1.463	2.830	0.8876	2.876	
$H_{KJ} \times 10^{-1}$	4.209	4.718	7.331	3.370	0.3229	
$H_{JK} \times 10^{-4}$	1.647	1.754	2.153	1.500	0.02490	
$H_J \times 10^{-9}$	-4.227	-4.295	-4.270	-4.181		
$h_J \times 10^{-9}$	-2.631	-2.695	-2.857	-2.548		
$h_{JK} \times 10^{-11}$	-2.213	-3.404	-10.68	1.201		
$h_K \times 10^{-10}$	3.556	3.719	4.567	3.196		

Theoretical values at CISD level of theory are taken from [128]

<sup>a</sup>Experimental effective rotational constants and quartic centrifugal distortion constants obtained from [129]

<sup>b</sup>Experimental electric dipole moment from [130]

Fig. 5 Rotational spectrum of HNCO at 300 K simulated with mPW2PLYPD/aug-cc-pVTZ level of theory (black). The experimental spectrum is shown in red



centrifugal distortion effects felt by this molecule. The simulated spectra at B2PLYPD/aug-cc-pVTZ and B97D/augcc-pVTZ levels of theory are shown in the SI.

The simulated and experimental spectra follow the same profile throughout all values of J. However, due to the unusually large error associated with the predicted  $A_0$  constant, calculated frequencies were significantly displaced in comparison with measured transitions. Overall, the mPW2PLYPD functional derived frequencies with discrepancies two orders of magnitude smaller than the ones obtained with B2PLYPD and B97D. More refined dispersion correction schemes and larger basis sets are likely to substantially improve those predictions.

#### Conclusions

In the present work, we briefly reviewed the most common model Hamiltonians to calculate the rotational spectroscopic constants of asymmetric tops beyond the rigid-rotor approximation. Those constants were used in the simulation of rotational spectra of some astrochemically relevant molecules. Spectroscopic constants calculated at DFT level of theory are overall close to measurements, resulting in rotational frequency predictions suitable for guiding laboratory experiments. Among the three chosen functionals, mPW2PLYP with dispersion corrections is the most successful in deriving effective rotational constants similar to measured values. In general, the GGA functional B97 with dispersion corrections predicted the most accurate quartic centrifugal distortion constants. However, its effective rotational constants are the furthest from experiment, resulting in the largest discrepancies in frequency predictions. Additionally, the  $B_0$  and  $C_0$  rotational constants have consistently been noted to play a dominant role in the simulations, being much more influential to the frequency predictions than the  $A_0$  constant for both  $C_{2v}$  and  $C_s$  symmetry species.

Moreover, the theoretical description of the  $\delta_K$  centrifugal distortion constant is likely to be neglecting important higher order terms, since its measured values are inconsistent with the other quadratic force field constants for H<sub>2</sub>CO, H<sub>2</sub>CS, c-HCOOH, and t-HCOOH. In the case of HNCO, the particularly strong centrifugal distortion is not described accurately enough at all levels of theory, requiring more refined theoretical treatments.

To the best of our knowledge, theoretical calculations of the centrifugal distortion constants of HNCO and c-HCOOH, as well as the effective rotational constants of c-HCOOH presented in this work, are being reported for the first time. The simulated spectra for the molecules considered here show a general good agreement with the experimental values, especially for transitions with J values up to 10, although the intensities are not so well reproduced in some cases.

Acknowledgements The authors acknowledge the National Laboratory for Scientific Computing (LNCC/MCTI, Brazil) for providing HPC resources of the SDumont supercomputer, which have contributed to the research results reported within this paper, and Fabio Cafardo for the help with coding and the dissertation of this text.

**Funding** The authors received financial support from Coordenação de Aperfeiçoamento de Pessoal de Nível Superior (CAPES) for financial support. Conselho Nacional de Desenvolvimento Científico e Tecnológico (CNPq).

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